

# Searching for frequency multiplets in the pulsating subdwarf B star PG 1219+534

**Missouri State**<sup>™</sup>  
U N I V E R S I T Y

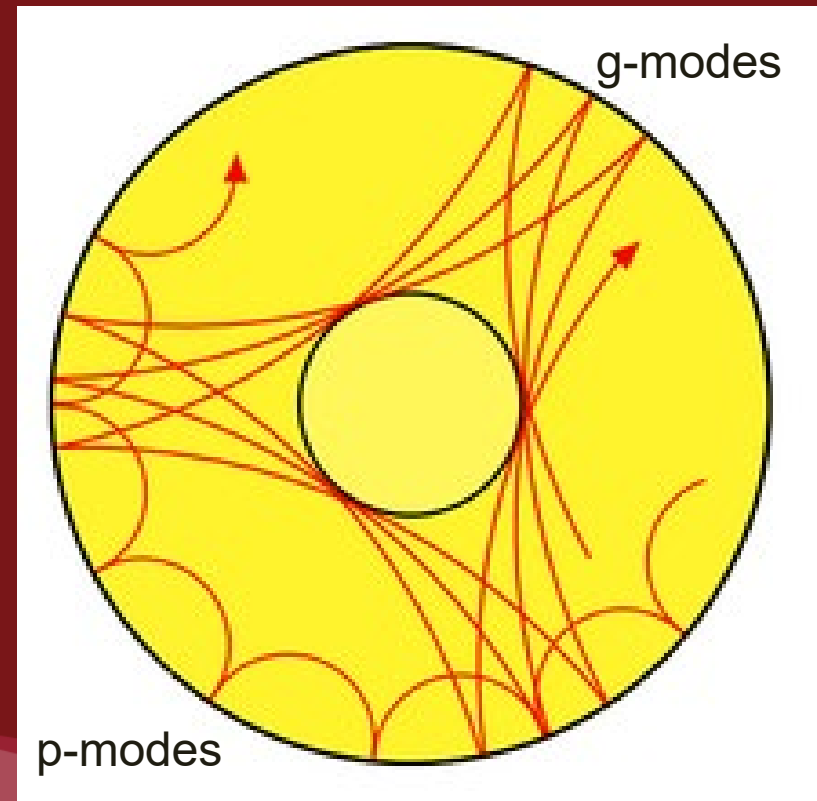
John Crooke, Ryan Roessler

Advisor: Mike Reed

Physics, Astronomy, and Materials Science

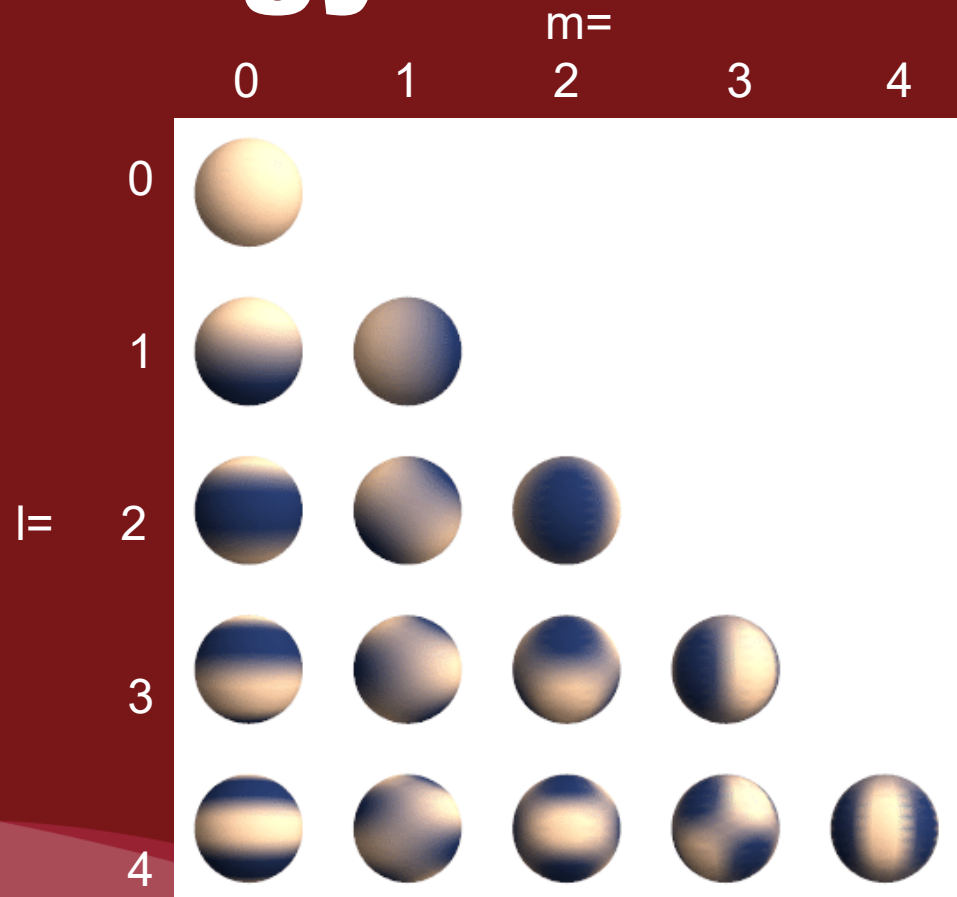
# Asteroseismology

- Study of pulsations to infer internal structure
  - Only way to “observe” the inside of a star
- Two oscillation modes
  - p: pressure mode, high frequency
  - g: gravity mode, low frequency



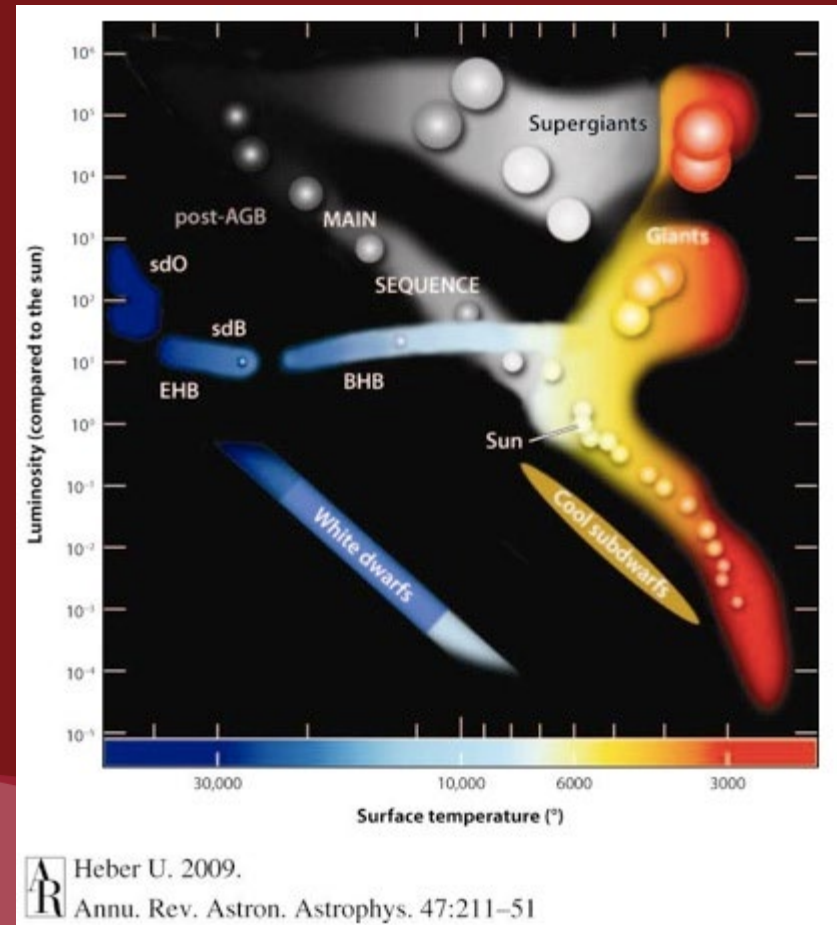
# Asteroseismology

- Using frequency multiplets (rotation period), it is possible to remove azimuthal degeneracy and associate pulsations with spherical harmonics
- Angular degree:  $l$ 
  - # of nodal lines
- Azimuthal order:  $m$ 
  - # of nodal lines crossing equator
- Radial order:  $n$ 
  - # of nodal lines along radius



# Subdwarf B (sdB) stars

- $0.5 M_{\text{sun}}, 0.2 R_{\text{sun}}$
- 20,000 – 30,000 K
- Horizontal Branch
  - post helium-flash
- Thin, inert envelopes<sup>1</sup>
- Helium fusing cores

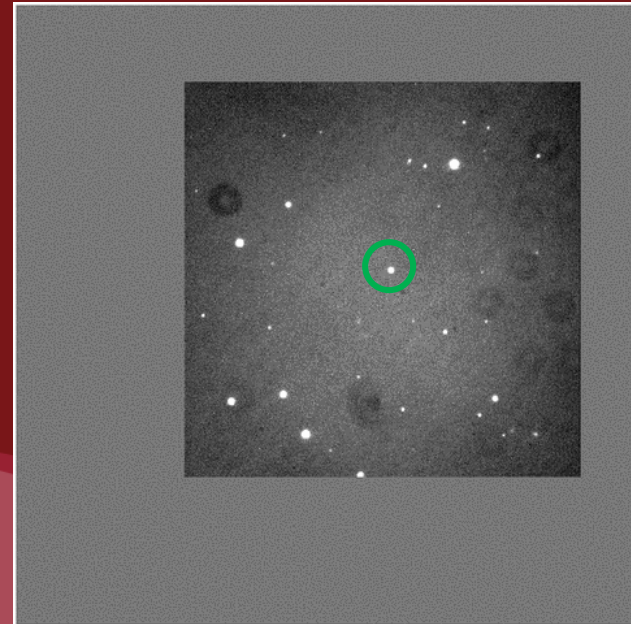


# Previous Work on PG1219+534

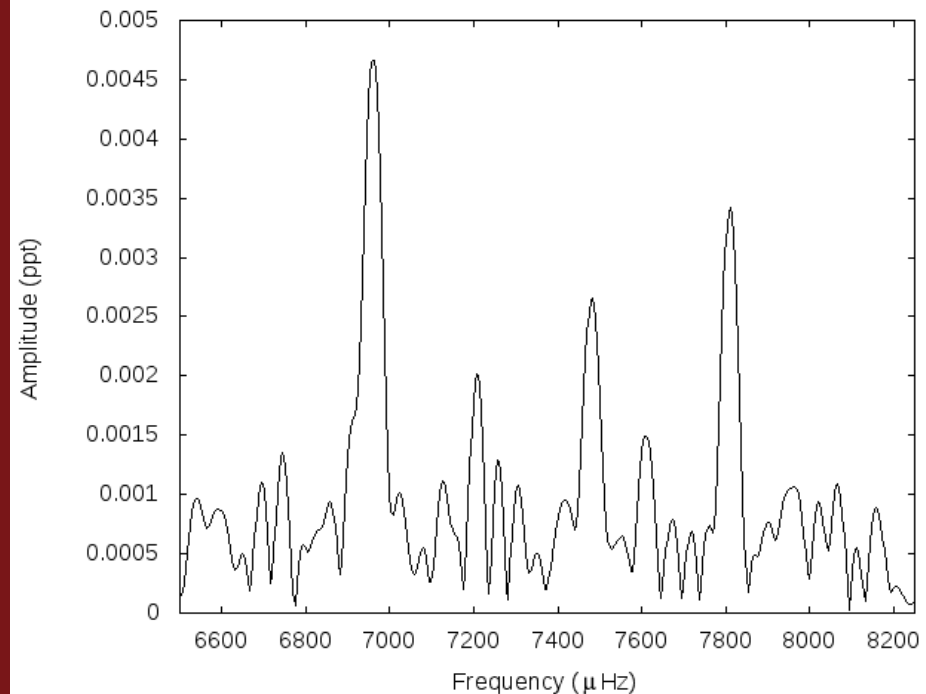
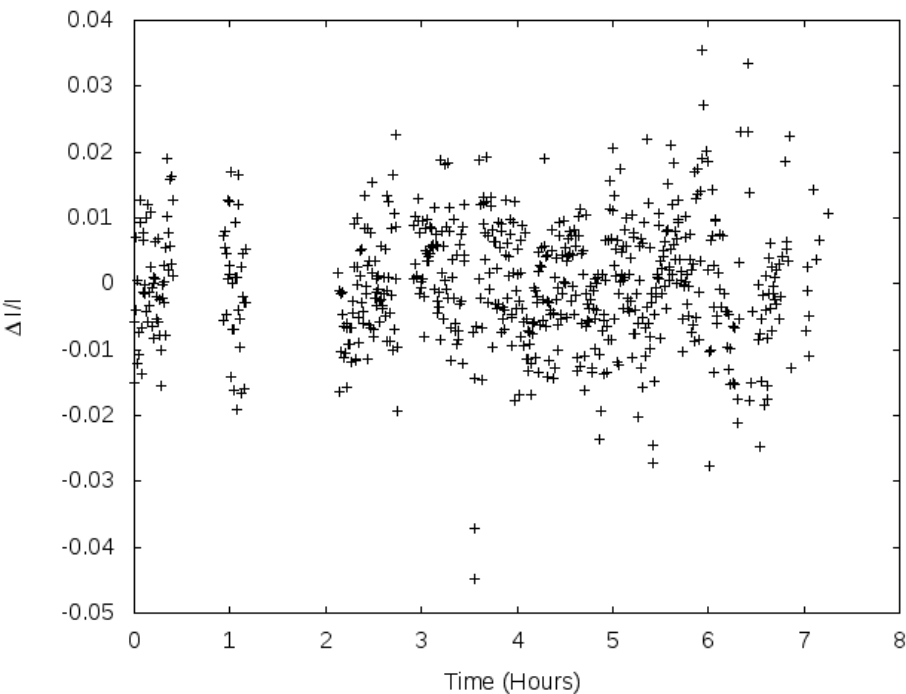
- Four dominant pulsation frequencies<sup>2</sup>
  - 6721  $\mu\text{HZ}$ , 148.8 seconds
  - 6961  $\mu\text{HZ}$ , 143.7 seconds
  - 7590  $\mu\text{HZ}$ , 131.8 seconds
  - 7807  $\mu\text{HZ}$ , 128.1 seconds
- Frequency multiplets indicating rotation period of  $\sim 35$  days<sup>3</sup>
  - No evidence or data published
- Goal of project is to find evidence of this rotation

# Observations

- January-August 2015, Baker Observatory
  - Baker Observatory Robotic Autonomous Telescope (BORAT)
    - 16-inch Schmidt-Cassegrain with Apogee U47 CCD
- 44 nights processed (28 usable), 125 hours of images total

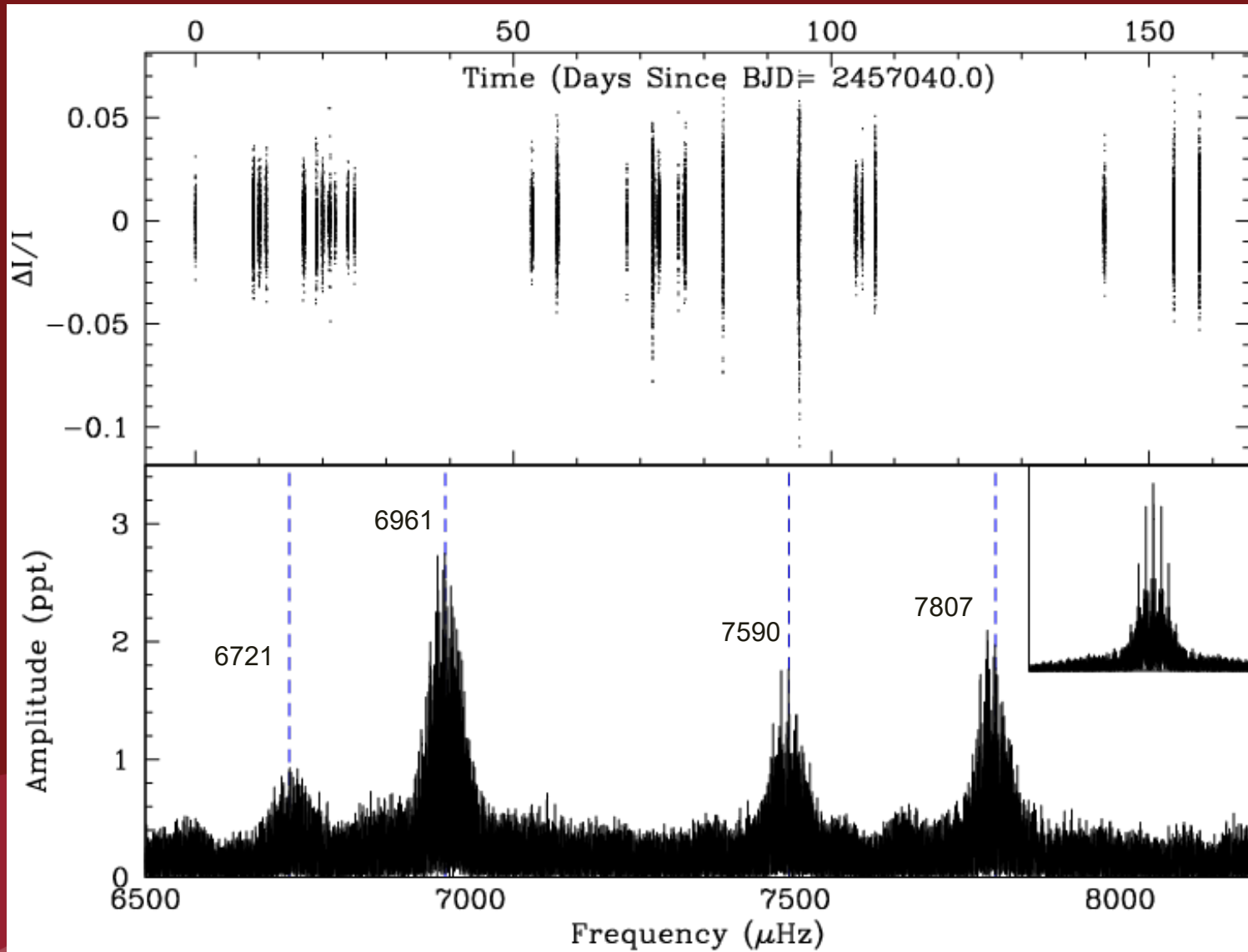


# Analysis



Sample lightcurve and Fourier Transform from March 31, 2015

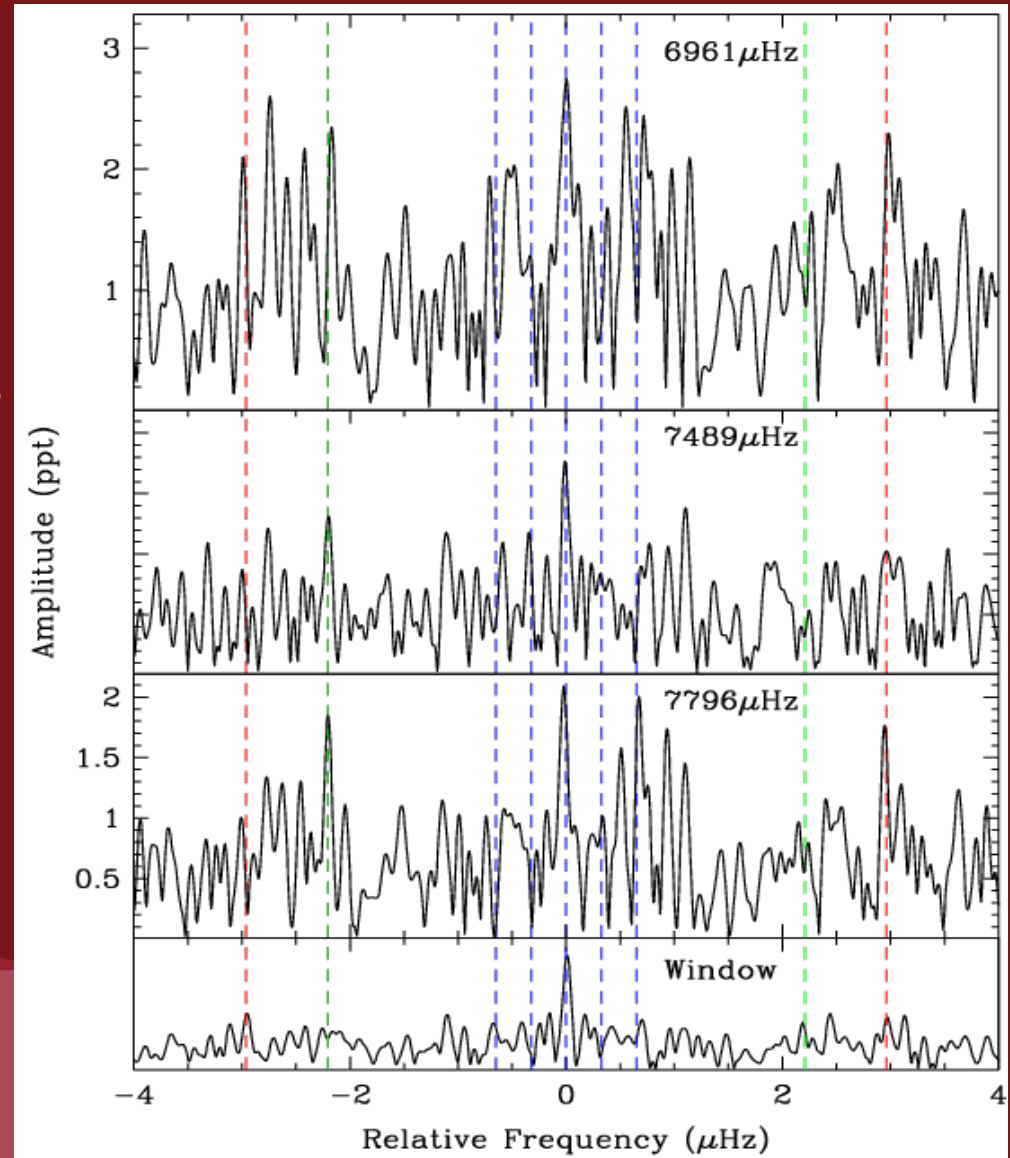
# Analysis





# Analysis

- Close-up view of dominant frequencies
- Blue lines have 0.3  $\mu\text{Hz}$  separation
  - Not consistent
- Green and red show that structure is not evenly split



# Conclusions

- Recovered three of four reported frequencies, fourth at low amplitude
- Do not see frequency multiplets indicating reported rotation period of 35 days
  - Should see it, data spans 160 days
- Have shown BORAT works for asteroseismology but with data density issues

# References / Acknowledgements

1. Heber, U., 2016, PASP, 128, 2001.
2. Reed, M., et al., 2013, A&A.
3. Fontaine, G., et al., 2014 ASPC, 481, 19.

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